

**\* PROBLEM 19: THE FREEZE-OUT OF A FICTITIOUS PARTICLE X**  
(25 points)

The following problem was Problem 3 of Quiz 3, 2016.

Suppose that, in addition to the particles that are known to exist, there also existed a family of three spin-1 particles,  $X^+$ ,  $X^-$ , and  $X^0$ , all with masses  $0.511 \text{ MeV}/c^2$ , exactly the same as the electron. The  $X^-$  is the antiparticle of the  $X^+$ , and the  $X^0$  is its own antiparticle. Since the  $X$ 's are spin-1 particles with nonzero mass, each particle has three spin states.

The  $X$ 's do not interact with neutrinos any more strongly than the electrons and positrons do, so when the  $X$ 's freeze out, all of their energy and entropy are given to the photons, just like the electron-positron pairs.

- (a) (5 points) In thermal equilibrium when  $kT \gg 0.511 \text{ MeV}/c^2$ , what is the total energy density of the  $X^+$ ,  $X^-$ , and  $X^0$  particles?
- (b) (5 points) In thermal equilibrium when  $kT \gg 0.511 \text{ MeV}/c^2$ , what is the total number density of the  $X^+$ ,  $X^-$ , and  $X^0$  particles?
- (c) (10 points) The  $X$  particles and the electron-positron pairs freeze out of the thermal equilibrium radiation at the same time, as  $kT$  decreases from values large compared to  $0.511 \text{ MeV}/c^2$  to values that are small compared to it. If the  $X$ 's, electron-positron pairs, photons, and neutrinos were all in thermal equilibrium before this freeze-out, what will be the ratio  $T_\nu/T_\gamma$ , the ratio of the neutrino temperature to the photon temperature, after the freeze-out?
- (d) (5 points) If the mass of the  $X$ 's was, for example,  $0.100 \text{ MeV}/c^2$ , so that the electron-positron pairs froze out first, and then the  $X$ 's froze out, would the final ratio  $T_\nu/T_\gamma$  be higher, lower, or the same as the answer to part (c)? Explain your answer in a sentence or two.

(a)

$$u = g \frac{\pi^2}{30} \frac{(kT)^4}{(\hbar c)^3}$$

$$g \begin{cases} 1 & \text{per spin state for bosons} \\ 7/8 & \text{per " " " fermions} \end{cases}$$

$X^+, X^-, X^0$

$$g = \underset{\substack{\uparrow \\ \text{\# of species}}}{3} \times \underset{\substack{\uparrow \\ \text{spin states}}}{3} \times \underset{\substack{\uparrow \\ \text{boson}}}{1} = 9$$

$$\Rightarrow u = \frac{3\pi^2}{10} \frac{(kT)^4}{(\hbar c)^3}$$

$$e^+e^-: g = \underset{\substack{\uparrow \uparrow \\ 1 \quad 1}}{(2)} \times \underset{\substack{\uparrow \\ \text{spin states}}}{2} \times \underset{\substack{\uparrow \\ \text{fermions}}}{7/8}$$

(b)

$$n = g^* \frac{\zeta(3)}{\pi^2} \frac{(kT)^3}{(\hbar c)^3}$$

$$g^* \begin{cases} 1 & \text{per spin state for bosons} \\ 3/4 & \text{" " " fermions} \end{cases}$$

$X^+, X^-, X^0$

$$g_{\text{tot}}^* = \underset{\substack{\uparrow \\ \text{species}}}{3} \times \underset{\substack{\uparrow \\ \text{spin states}}}{3} \times \underset{\substack{\uparrow \\ \text{boson}}}{1}$$

$$\Rightarrow n = \frac{9 \zeta(3)}{\pi^2} \frac{(kT)^3}{(\hbar c)^3}$$

(c) After neutrinos decouple, but before  $X$ ,  $e$  freeze out,

$$S_{\text{tot}} = S_\nu + S_{\text{other}} \Rightarrow a^3 S_\nu \text{ is constant}$$

separately conserved

$$a^3 S_{\text{other}} \text{ is constant}$$

In general,  $S = A g T^3 \Rightarrow g_\nu(a_\pm^3 T^3)_{\text{after neutrinos decouple}} = g_\nu(a_F^3 T_\nu^3)_{\text{any later time}}$

$$(g_{\text{other}} a_\pm^3 T^3)_{\text{after neutrinos decouple}} = (g_{\text{other}} a_F^3 T_\gamma^3)_{\text{after electrons, } X \text{ freeze out}} \quad (*)$$

$$g_{\text{other}}^{\text{before freezeout}} = \begin{matrix} e^+e^- \\ 2 \times 2 \times 7/8 \end{matrix} + \begin{matrix} \gamma \\ 1 \times 2 \times 1 \end{matrix} + \begin{matrix} X^{+/-, 0} \\ 3 \times 3 \times 1 \end{matrix}$$

species      spin states      fermions      species      polarizations      bosons      species      spin states      bosons

$$= 7/2 + 2 + 9 = 29/2$$

$$g_{\text{other}}^{\text{after freezeout}} = 2$$

$$(*) \Rightarrow \frac{29}{2} a_\pm^3 T^3 = 2 a_F^3 T_\gamma^3$$

$S_\nu$  conservation

$$= a_F^3 T_\nu^3$$

$$\Rightarrow \frac{29}{2} \cancel{a_F^3} T_\nu^3 = 2 \cancel{a_F^3} T_\gamma^3 \Rightarrow T_\nu = \left(\frac{4}{29}\right)^{1/3} T_\gamma$$

(d)  $\frac{T_\nu}{T_\gamma}$  is exactly the same as in part (c).

The entropy must be conserved from the initial state (with  $\gamma, e, X$ ) to the final state (with only  $\gamma$ ).